## **Gravitational Waves**

Physics, Technology, Astronomy and Opportunities

Unnikrishnan. C. S. Tata Institute of Fundamental Research Mumbai

> unni@tifr.res.in www.tifr.res.in/~filab





Four Fundamental Interactions and several fundamental particles

Gravity, Electromagnetism, Weak interactions and sub-nuclear strong interactions.

Electromagnetism Weak Interaction Strong Interaction







**Electric Field** 

Magnetic Field

Charges and Currents are fundamental, fields are 'theoretical'.

Current  $\rightarrow$  Magnetic field

Electromagnetic waves need electric AND magnetic field for generation and propagation.

$$-\frac{\partial B}{\partial t} = \nabla \times E, \quad \mu_0 \varepsilon_0 \frac{\partial E}{\partial t} = \nabla \times B \quad \dots Maxwell$$

**Unobservables in Physics** 

Fields, Wavefunction, Space and Time...



### GRAVITY, ITS FIELDS and THEIR WAVES





$$F = kQq/r^{2}$$

$$\vec{a} = \vec{F}/m$$

$$a_{em} = kQq/r^{2}m = \frac{kQ}{r^{2}}\left(\frac{q}{m}\right)$$

Gravity seems to be a special interaction

Inertia turns out to be identical to the gravitational charge – Equivalence Principle (physics of gravity identical to physics in accelerated frames)



### Universality of Free-Fall



$$\delta a < 10^{-12} \, m \, / \, s^2 \rightarrow \frac{\delta a}{a} < 10^{-13}$$



So, gravitational field 'g' and acceleration '-a=g' seem equivalent

## This is called the Equivalence Principle

This is the same as saying that in free-fall, there is no gravitational field

But it does not mean that in free-fall there is no gravity!





So, tidal deviations cannot be eliminated by free-fall (description in the General Theory of Relativity)



### Torsion balance: Harmonic potential Ultra-sensitive







*Flux – density* ∝ 1 / *R* (For 3D space)

What if space is higher dimensional at some tiny scales (micrometers or less)?

Inverse-square law for 'g' will change!











ravitational Waves

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'g' of Sun ~0.6 cm/s<sup>2</sup>

### $\delta a(earth-moon) < 10^{-14} cm/s^2$





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# 3% of total estimated matter: So 97% is invisible, and NOT made of ANY known particle

So, we are not sure of gravity's behaviour at very small scales and at very large scales.





Current  $\rightarrow$  Magnetic field

Electromagnetic waves need electric AND magnetic field for generation and propagation.

$$-\frac{\partial B}{\partial t} = \nabla \times E, \quad \mu_0 \varepsilon_0 \frac{\partial E}{\partial t} = \nabla \times B \quad \dots Maxwell$$



### What is an electromagnetic wave?



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### Gamma rays and galaxies









### Spectral view – multi-wavelength





### Spectral view – multi-wavelength

OPTICAL X-RAY . RADIO INFRARED

Crab Nebula



#### Multi-wavelength galaxy



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### Gravity and Electromagnetism

Both have 'electric' and 'magnetic' parts  $\rightarrow$  Charges and currents.

Mass is the 'charge' of gravity and Spin is its 'gravito-magnetic moment'

One important different between the two is that while electric and magnetic fields have no electric charge, gravitational field has gravitational charge!

With  $m=E/c^2$ , all forms of energy is equivalent to mass, and hence generate gravity. Therefore, all fields including the gravitational field, which carry energy, also generate gravitational fields. This is one reason why the theory of gravity (The General Theory of Relativity) is complicated to work with.



### Gravity and electromagnetism

Charges (static): Coulomb force – electric fields – Electromagnetic Currents (motion): Ampere's force – magnetic fields ✓ Waves

What about relativistic gravity?

We know static gravitational charge (mass/energy) generates g-field. Does moving and rotating masses generate a gravito-magnetic field?!

If so, then there is a possibility of gravitational waves...



### Gravito-magnetism

A natural consequence of relativistic gravity, and yet, was not detected experimentally till recently.



$$B = \frac{\mu_0 M}{r^3}$$



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### The real gravitational field near the earth





$$\omega \approx \frac{GM}{c^2 r} \Omega \approx 5 \times 10^{-14} rad / s$$

**Gravity Probe – B (Stanford U.)** 



### **Physics of Gravitational Waves**

What is an electromagnetic wave?





So, can we expect gravitational waves from oscillating (accelerated) masses?





#### Radiation

$$\frac{E_t}{E_r} = \frac{vt}{ct} \approx \frac{at}{c} = \frac{aR}{c^2} \text{ and } E_r = q/R^2$$
$$E_t = \frac{q}{R^2} \frac{aR\sin\theta}{c^2} = \frac{q\ddot{r}\sin\theta}{Rc^2}$$
$$qr = d \longrightarrow E_t = \frac{\ddot{d}\sin\theta}{Rc^2}$$



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Since area=R<sup>2</sup>, number of flux lines/area  $\rightarrow 1/R^2$ So, radial Electric field (flux/area) ~ 1/R<sup>2</sup>

What about the Transverse Electric field (radiation)?

Since, circumference of a great circle on the sphere increases only as R, transverse radiation field decrease as 1/R.



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### The force-field of gravitational waves











#### The relation between spin of the field and polarization of the force field



#### What is the physical effect of a passing gravitational wave?



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## Are we confident that Gravitational Waves exist, apart from the belief in the correctness of the theory?

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### ON GRAVITATIONAL WAVES.

BY

Jl. Franklin Inst. 1937

#### A. EINSTEIN and N. ROSEN.

#### ABSTRACT.

The rigorous solution for cylindrical gravitational waves is given. For the convenience of the reader the theory of gravitational waves and their production, already known in principle, is given in the first part of this paper. After encountering relationships which cast doubt on the existence of *rigorous* solutions for undulatory gravitational fields, we investigate rigorously the case of cylindrical gravitational waves. It turns out that rigorous solutions exist and that the problem reduces to the usual cylindrical waves in euclidean space.

Note.—The second part of this paper was considerably altered by me after the departure of Mr. Rosen for Russia since we had originally interpreted our formula results erroneously. I wish to thank my colleague Professor Robertson for his friendly assistance in the clarification of the original error. I thank also Mr. Hoffmann for kind assistance in translation.

> A. EINSTEIN. Gravitational Waves

### Binary Pulsar 1913+16 (Hulse-Taylor)



Binary pulsar



 $t_c \approx 3 \times 10^8 \text{ yrs}$ 

$$\dot{E}_G = \frac{dE_G}{dt} \approx \frac{32G}{c^5} M^2 \omega^6 r^4$$





 $\omega^2 \approx GM / a^3 \rightarrow \nu \approx \frac{1}{2\pi} \sqrt{G\rho}$ 

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Orbital decay and speeding up of the binary pulsar:



### Signal Strength at Earth for neutron star spiral in milky way:

Distance: 10 kpc ~ 10<sup>20</sup> meters

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Strain 
$$h \approx \frac{G}{c^4} \frac{Mv^2}{R} \approx 10^{-64} Mv^2$$

 $v \approx \sqrt{\frac{Gm}{2r}} = 0.1c \ (3 \times 10^7 \text{ m/s}) \text{ for neutron stars at } r \approx 100 \text{ km}$ 

With M~10<sup>30</sup> kg, v~3x10<sup>7</sup> m/s, Strain 
$$h = \frac{\Delta l}{l} \approx 10^{-19}$$

If the event happens in another galaxy, 100 Mpc ( $10^{24}$  m away), Strain  $h \approx 10^{-23}$ 

This small strain requires the measurement of  $<10^{-20}$  meters in a detector of size 1 km! (almost million times smaller than the atomic nucleus).

### Is it a mad venture trying to make a 'detector'?



#### EVIDENCE FOR DISCOVERY OF GRAVITATIONAL RADIATION\*

#### J. Weber

Department of Physics and Astronomy, University of Maryland, College Park, Maryland 20742 (Received 29 April 1969)

Coincidences have been observed on gravitational-radiation detectors over a base line of about 1000 km at Argonne National Laboratory and at the University of Maryland. The probability that all of these coincidences were accidental is incredibly small. Experiments imply that electromagnetic and seismic effects can be ruled out with a high level of confidence. These data are consistent with the conclusion that the detectors are being excited by gravitational radiation.



### $h \approx 10^{-17}$
A modern cryogenic resonant mechanical detector



When these waves reach earth, what can they do to free masses?



Strain 
$$h \approx \frac{G}{c^2 R} \frac{M v^2}{c^2} \rightarrow \frac{6.7 \cdot 10^{-11} \times 3 \cdot 10^{30}}{(3 \times 10^8)^2 \times 10^{20}} 0.05^2 < 10^{-19} m / m!$$

Much less than the size of the nucleus. This is the primary device for gravitational wave detection





Thermal noise, Seismic Noise, Quantum Noise Tidal Noise, All instrument Noise, Any Noise one can think of...

Signal 1/R, G/c<sup>2</sup>, Random source..



Lower Limits of all these noises allowed by Physics and today's technology...





With 1 W of optical power,  $N=10^{19} / s$ ,

$$\Delta \phi_{\min} \approx \frac{\pi \Delta N_{\min}}{N} = \frac{\pi}{\sqrt{N}} \approx 10^{-9}$$



There is another equivalent way to talk about photon shot noise that explicitly brings out the basic feature of quantum mechanics involved.

The energy-time uncertainty relation  $\Delta E \Delta t \geq \hbar$ 

 $\Delta E \Delta t = \hbar \omega (\Delta N) \Delta \phi / \omega \approx \hbar$  $\rightarrow \Delta N \Delta \phi \ge 1 \rightarrow \Delta \phi \approx \frac{1}{\Delta N} = \frac{1}{\sqrt{N}}$ 



Detection of gravitational waves requires the measurement of movements 10<sup>-17</sup> to 10<sup>-20</sup> meters in a detector of size 1 km.



2) Increase Length up to 4 km: Not much gain, though very important (1/R).



Reaching there, but not comfortable yet!

Optimal length of the Interferometer arm:

$$L_{opt} \approx \frac{\lambda_g}{4} \approx \frac{c / v_g}{4} \approx \frac{3 \times 10^8 \, m / 100 Hz}{4} = 750 \, km!$$

This is the optimal distance the light should travel for maximal signal. In other words light should be in the interferometer for an optimal duration of about 750 km/c seconds or a quarter of the GW period of 10 ms or so. This is achieved by multiple bounces with average time equal to about quarter of the GW period such that 4 km x  $n_B = 750$  km. So, the number of bounces is about 200.

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Detection of gravitational waves requires the measurement of movements 10<sup>-17</sup> to 10<sup>-19</sup> meters in a detector of size several km. Improvements:

1) Folding  $\rightarrow$  Fabry-Perot Cavity



10 kW, 4 km  $\rightarrow$  10<sup>-16</sup> m  $\rightarrow$  3×10<sup>-19</sup> m with F  $\approx$  300

Intra-cavity power > 1 MW !

Radiation Pressure Noise and Thermal Lensing are problems

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### Large Interferometer VIRGO at Pisa, Italy (3 km)







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LIGO-HO



Strain 
$$h = \frac{\Delta L}{L} \approx \frac{G}{c^4} \frac{Mr^2 \omega^2}{R}$$

$$\omega^2 \approx GM / a^3 \rightarrow v \approx \frac{1}{2\pi} \sqrt{G\rho}$$



## But, every bit counts... because waves strength is 1/R

If sensitivity is increased by factor X, then the distance reach increases by X, and the number of astrophysical sources increases as X<sup>3</sup>!

So, a factor of 10 in sensitivity means a factor of 1000 in number of possible detections.



Scheme of the Interferometric detector

3m scale prototype being built in TIFR – can measure <10<sup>-17</sup> m



# **Schematic of Advanced LIGO detectors**



10 mega -litres at nano-torr

Noise from Light:

Photon Shot Noise:  $h_{st}$ 

$$\Delta l = \lambda \Delta \phi_{\min} = \frac{\lambda \pi}{\sqrt{N}}$$
$$= \alpha \sqrt{\frac{1}{P_i}} \qquad \qquad \sqrt{N} = \sqrt{\frac{P_i}{hc/\lambda}}$$

Radiation Pressure Noise  $F_{rad} = \sqrt{N} (h/\lambda) = \sqrt{\frac{hP_i}{c\lambda}}$ 

Movement noise due to this force:  $h_{rp} \propto F_{rad} / m\omega^2 = \beta \sqrt{P_i}$ 

Standard Quantum Limit

$$h_{total} = \sqrt{h_{sn}^2 + h_{rp}^2} \rightarrow h_{min} = h_{SQL} = \frac{1}{\pi L f} \sqrt{\frac{\hbar}{m}}$$



How can we hope to measure 10<sup>-19</sup> m when the ground vibrations are like 1 micrometer?!

Immunity to vibrations needed by a factor of 10<sup>14</sup>!



Ground vibrations: 10<sup>-6</sup> m at 1 Hz, 10<sup>-9</sup> m at 30 Hz

3 stages of springs and pendulum with each resonance at around 1 Hz  $\rightarrow$  Response down by a factor  $(10^3)^3 = 10^9$  at 30 Hz!

Possible to isolate from vibrations at the 10<sup>-20</sup> m level at 100 Hz with 3-4 stages. IIT-K, 2014 Gravitational Waves





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### Absorption/Scattering $\rightarrow$ ppm level



GEO600

$$h_{total} = \sqrt{h_{sn}^2 + h_{rp}^2} \rightarrow h_{min} = h_{SQL} = \frac{1}{\pi Lf} \sqrt{\frac{\hbar}{m}}$$



### A sample of Vacuum chambers (LIGO):

### < 10<sup>-8</sup> mbar, all 4x2 km





Quantum Noise

 $\Delta x \Delta p \ge \hbar$ Light: Energy  $\mathbf{E} \approx E^2 + B^2 \rightarrow p^2 + q^2$  Harmonic Oscillator

Main features: Zero point energy, equally quantized energy levels interpreted as number states of photons







Therefore, if we can somehow SQUEEZE the uncertainty is amplitude or phase, we can improve measurements by the squeezing factor.

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If the injected 'vacuum' is squeezed by a factor 10, then the quantum noise is reduced by the same factor, and this is equivalent to increasing the laser power by a factor 100, because noise reduces only as square root of power!





Roman Schnabel American JI. Physics 2013



Why do we need more detectors?

GW detectors are not telescopes. They cannot pin-point a source.



Timing (and only timing) can fix a direction

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Need at least 3

### The LIGO-India Concept



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Good angular resolution IIT-K, 2014

Base-lines to Japan comparable Gravitational Waves

#### The LIGO-India Idea and Opportunity

The LIGO-India proposal envisages the hardware meant for one of the LIGO detectors (out of 3) to be given to India and Indian scientists and engineers will build and operate the detector at a suitable site in India. Thus the LIGO-India detector will be the third vertex of the LIGO network, working like a large gravitational wave telescope.

The Dept. of Atomic Energy, Dept. of Science and Technology and the Planning Commission have approved the Rs.1300 Crore (15 years) proposal. A cabinet approval is required and awaited. Site selection is in progress.



www.gw-indigo.org





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### **Source localization error**

S. Fairhurst arXiv:1205.6611v2

B. S. Sathyaprakash et al., LIGO document T1200219-v1



### Original Plan 2 +1 LIGO USA+ Virgo



60° 75° AHLV 45° 30° -150°120°-90°-60°-30° 0° 30° 60° 90° 120°150° -15° -30° 45° -60° -75°

LIGO-India plan ျ<del>က</del>ြန် ၂၃၆၄စု USA+ Virgo+ LIGO-India LIGO-Aus plan 1+1 LIGO USA+ Virgo+LIGO-Aus Waves











# Gravitational Waves: The Indian Initiative LIGO-India Project

There are great possibilities and bright future for gravitational wave-based astronomy... if we manage to detect gravitational waves with these detectors

The IndiGO Consortium

www.gw-indigo.org



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